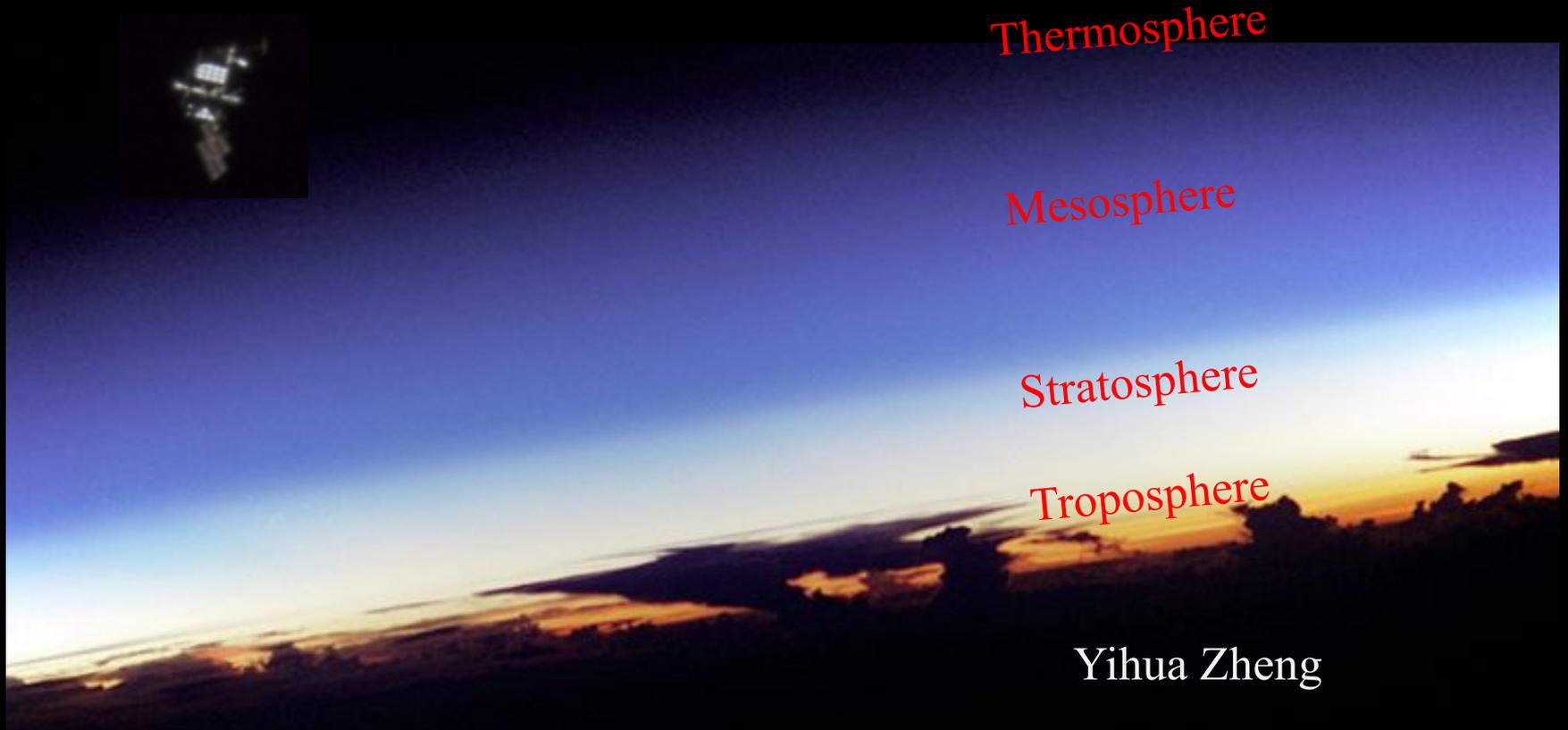


Satellite Drag



Thermosphere

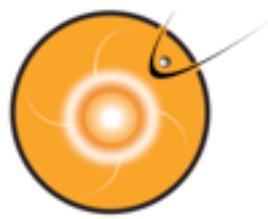
Mesosphere

Stratosphere

Troposphere

Yihua Zheng

Acknowledge: Prof . Delores Knipp
University of Colorado Boulder



Space Weather in the Thermosphere: Satellite Drag



Motivation:

- Track and identify active payloads and debris
 - Collision avoidance and re-entry prediction
 - Attitude Dynamics
 - Constellation control
 - “Drag Make-Up” maneuvers to keep satellite in control box
 - Delayed acquisition of SATCOM links for commanding /data transmission
 - Mission design and lifetime
-
- Study the atmosphere’s density (and temperature) profiles

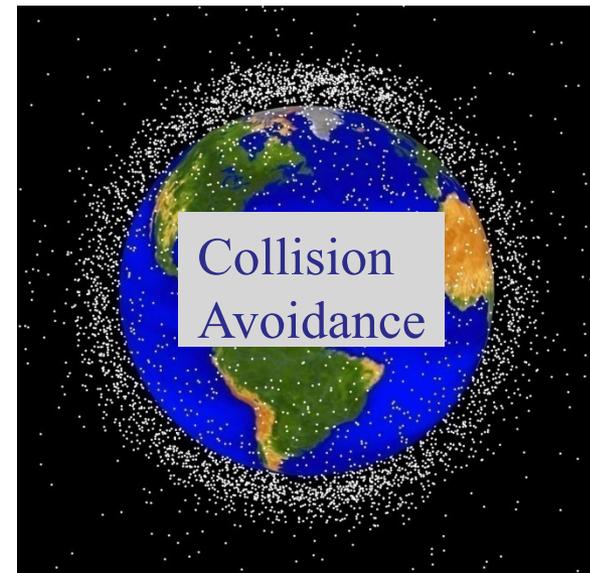
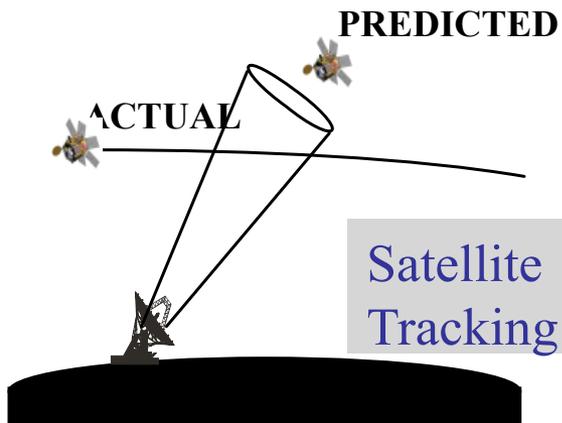
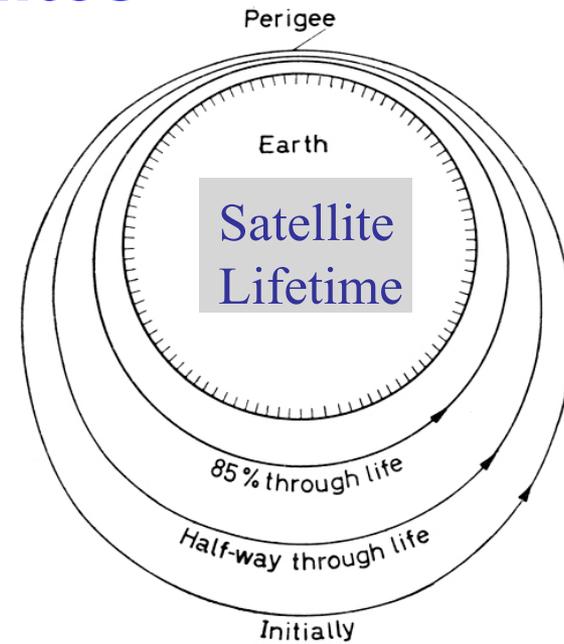
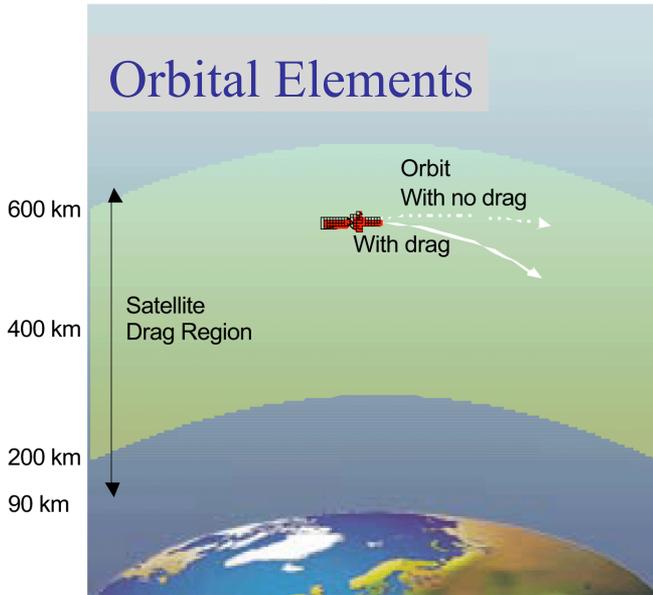
Overview

- Fundamentals of Satellite Drag
- Thermosphere and Its Characteristics
- Dynamic Space Weather Drivers
- Collision Avoidance

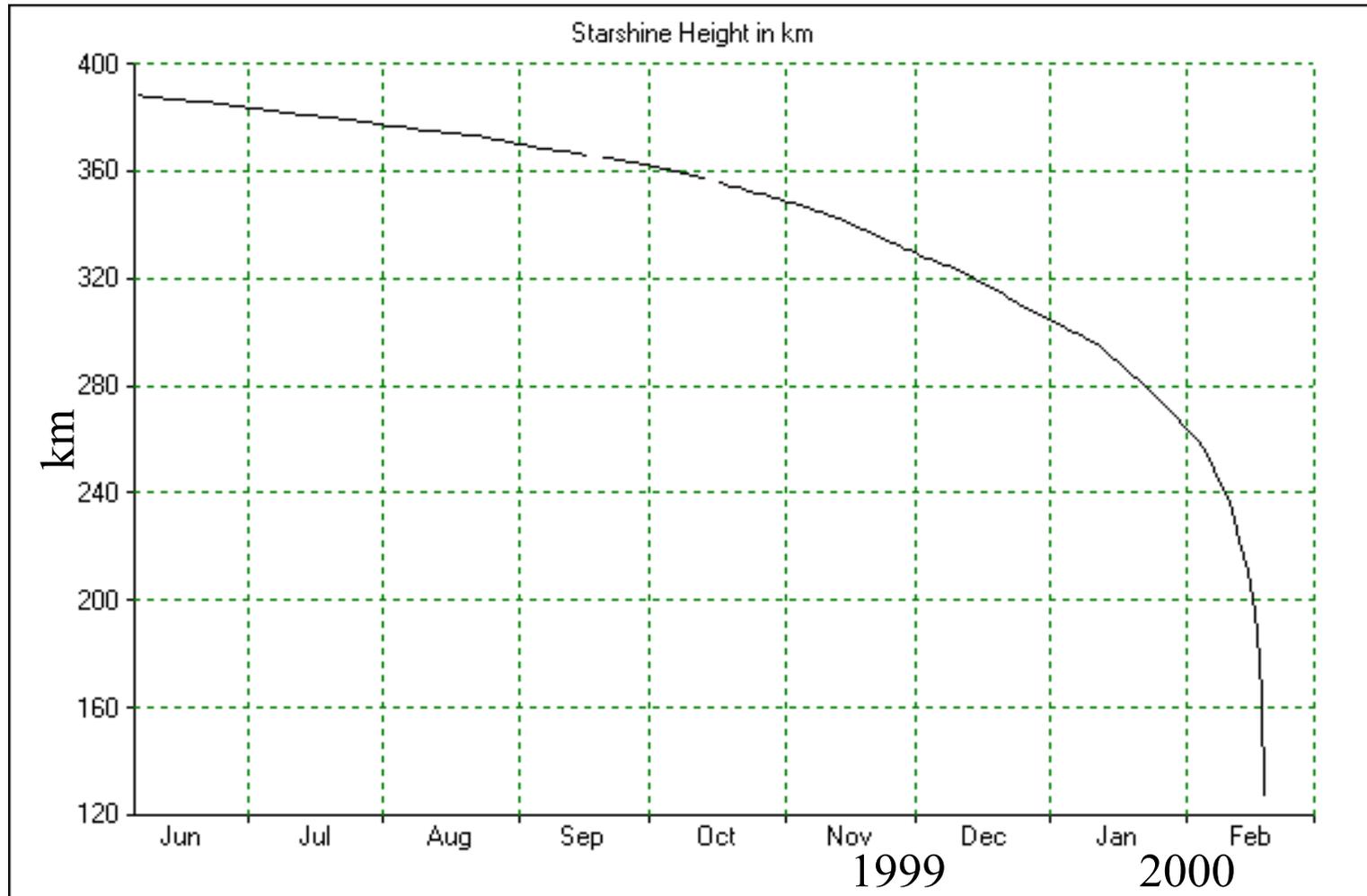
Spacecraft Drag

- Spacecraft in LEO experience periods of increased drag that causes them to speed-up, lose altitude and finally reenter the atmosphere. Short-term drag effects are generally felt by spacecraft <1,000 km altitude in an atmospheric region called the thermosphere.
- Drag increase is well correlated with solar Ultraviolet (UV) output, and atmospheric heating that occurs during geomagnetic storms. Recently lower atmospheric tidal effects have been modeled in satellite drag response.
- Most drag models use solar radio flux at 10.7 cm wavelength as a proxy for solar UV flux. Kp/Ap are the indices commonly used as a surrogate for short-term atmospheric heating due to geomagnetic storms. In general, 10.7 cm flux >250 solar flux units and Kp>=6 result in detectably increased drag on LEO spacecraft.
- Very high UV/10.7 cm flux and Kp/Ap values can result in extreme short-term increases in drag. During the great geomagnetic storm of 13-14 March 1989, tracking of thousands of space objects was lost. One LEO satellite lost over 30 kilometers of altitude, and hence significant lifetime, during this storm.

Atmospheric Drag on Satellites



STARSHINE-1 Height vs Time Profile



Satellite Drag and Thermosphere Density

Aerodynamic forces are the forces created by a spacecraft's movement through a neutral density atmosphere. The forces result from momentum exchange between the atmosphere and the spacecraft and can be decomposed into components of lift, drag, and side slip.

$$F_i = \frac{1}{2} \rho A C_i v^2$$

F_i = Force, N

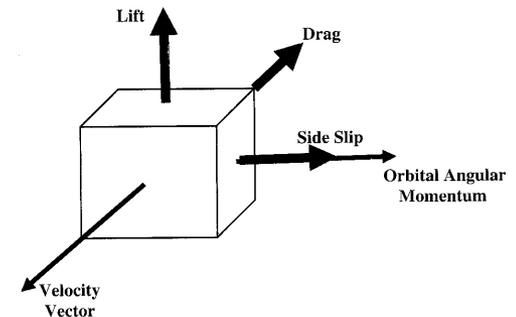
i = d(drag), l(lift), and s (side slip)

A = area, m^2

C_i = coefficient

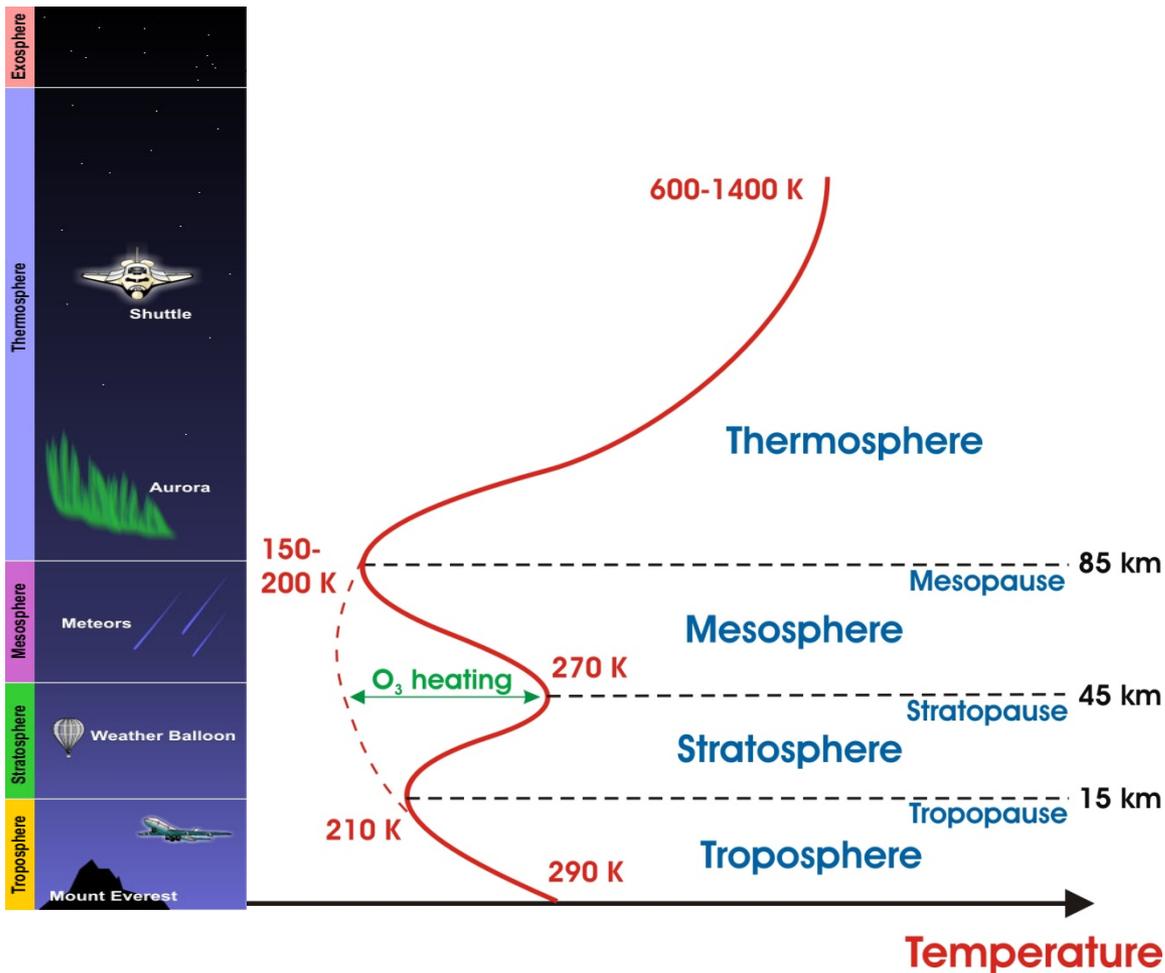
v = spacecraft speed with respect to the atmosphere, m / s

ρ = atmospheric mass density, kg / m^3



In the thermosphere the density is a function of temperature

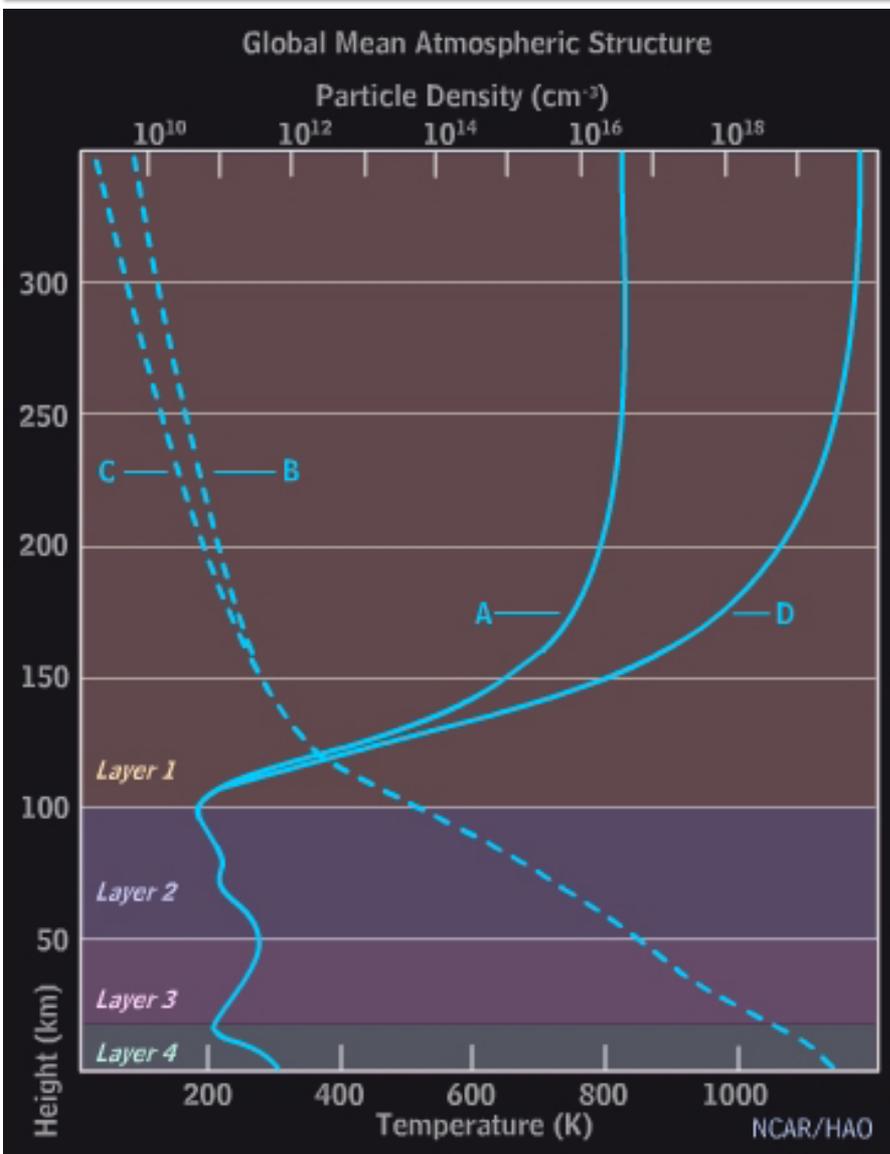
Upper Atmosphere – Thermosphere



The outer gaseous shell of a planet's atmosphere that exchanges energy with the space plasma environment: **Thermosphere**

- **Energy sources:**
 - Absorption of Extreme UV radiation (10-200nm)
 - Joule heating by electrical currents
 - Particle precipitation from the magnetosphere
 - Dissipation of upward propagating waves (tides, planetary waves, gravity waves)
- **Energy sinks:**
 - Thermal conduction into the mesosphere
 - IR cooling by CO₂, NO, O
 - Chemical reactions

Thermosphere Variability and Time Scales



Thermosphere:

• Characteristics

- Very high temperatures, often exceeding 1000 K
- Low neutral density
- Matter sorted by gravity—heavier material at base
- Dominated by atomic oxygen

• Time Scales

- Solar cycle
- Annual
- 27 day
- Equinoctial
- Day /night

A Solar Min Temperature
D Solar Max Temperature

C Solar Min Density
B Solar Max Density

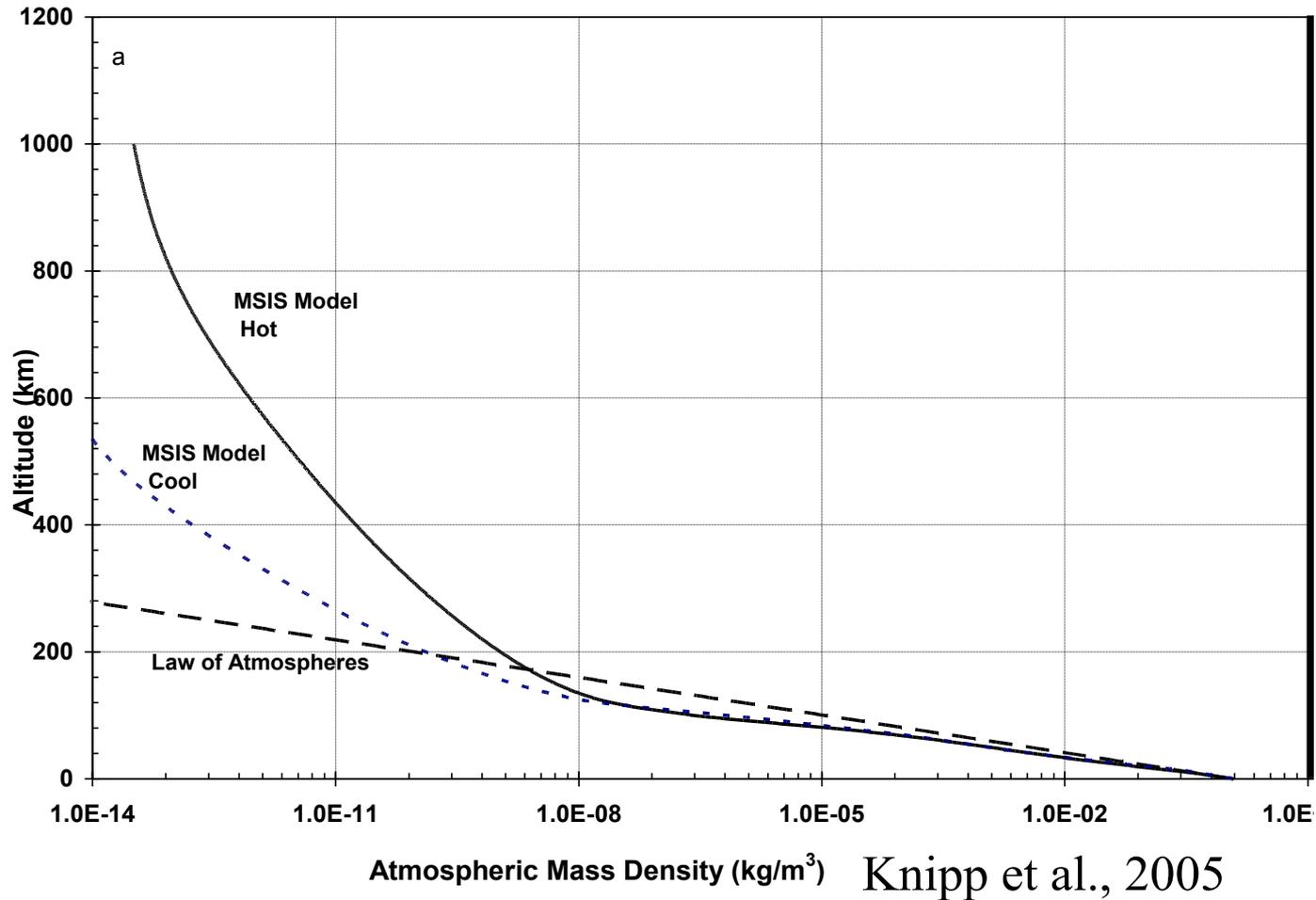
Courtesy of UCAR COMET

Density Variations at 400 km

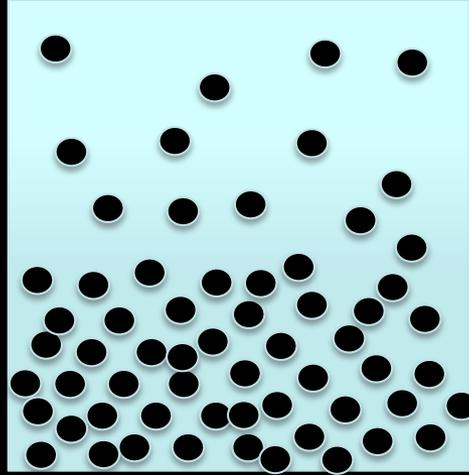
Variations	change	frequency
Solar cycle	1600%	11 years
Semiannual	125%	12 months
Solar UV rotation	250%	27 days
Major geomagnetic storm	800%	3 days
Diurnal effect	250%	1 day

Ideal and Model Atmosphere Neutral Density

Altitude vs. Atmospheric Mass Density, Comparing Different Models



Thermosphere



Exponential Atmosphere

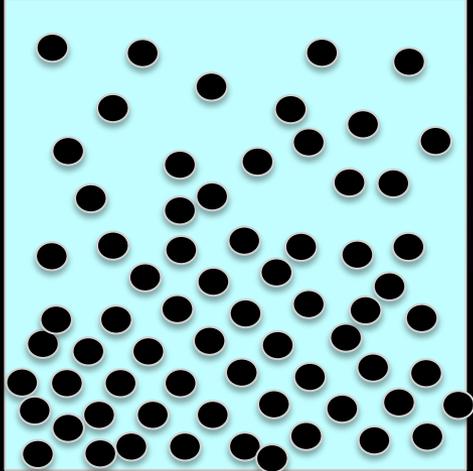
Mesosphere

Polar Mesospheric Clouds 80-85 km

Stratosphere

Troposphere

You are here ↑



Thermosphere

Hotter Exponential Atmosphere

Mesosphere

Polar Mesospheric Clouds 80-85 km

Stratosphere

Troposphere

You are here ↑

Satellite Drag and Thermosphere Density

The drag force is considered the most dominant force on low-earth orbiting spacecraft and serves to change the energy of the spacecraft through the work done by the drag force.

$$\frac{dE}{dt} = F_d \cdot v = -\frac{1}{2} \rho C_d A v^3$$

F_i = Force, N

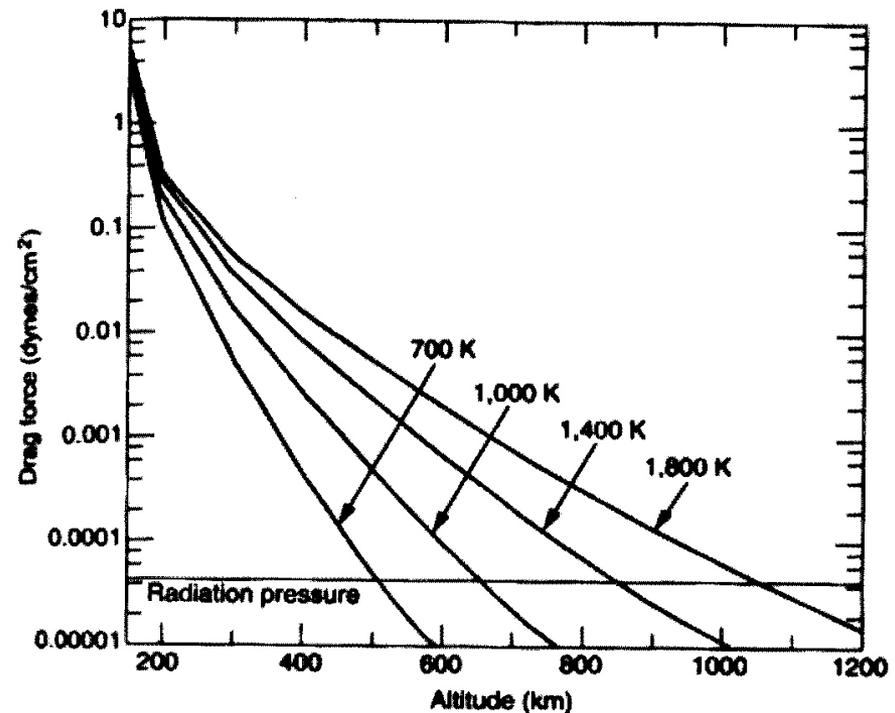
i = d(drag), l(lift), and s (side slip)

A = area, m^2

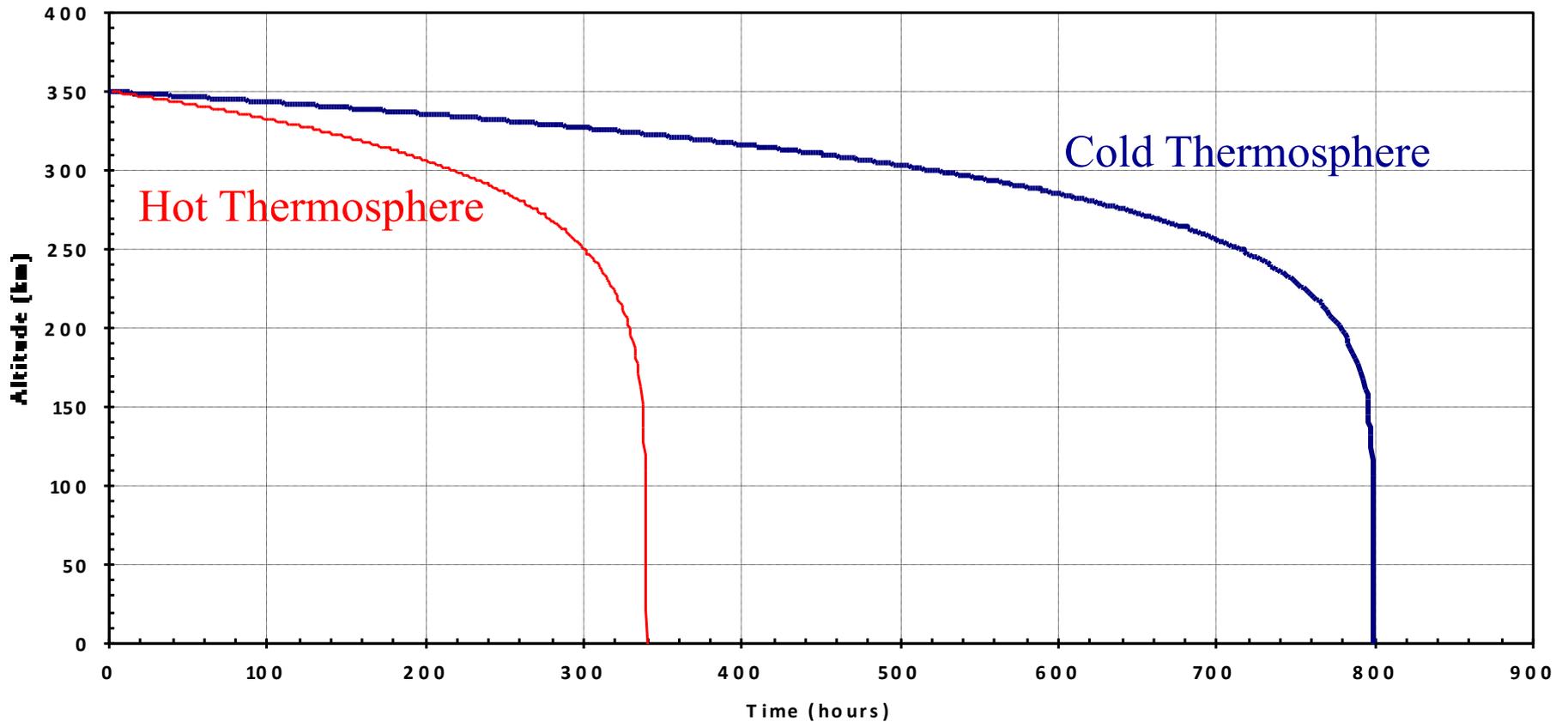
C_i = coefficient

v = spacecraft speed with respect to the atmosphere, m / s

ρ = atmospheric mass density, kg / m^3

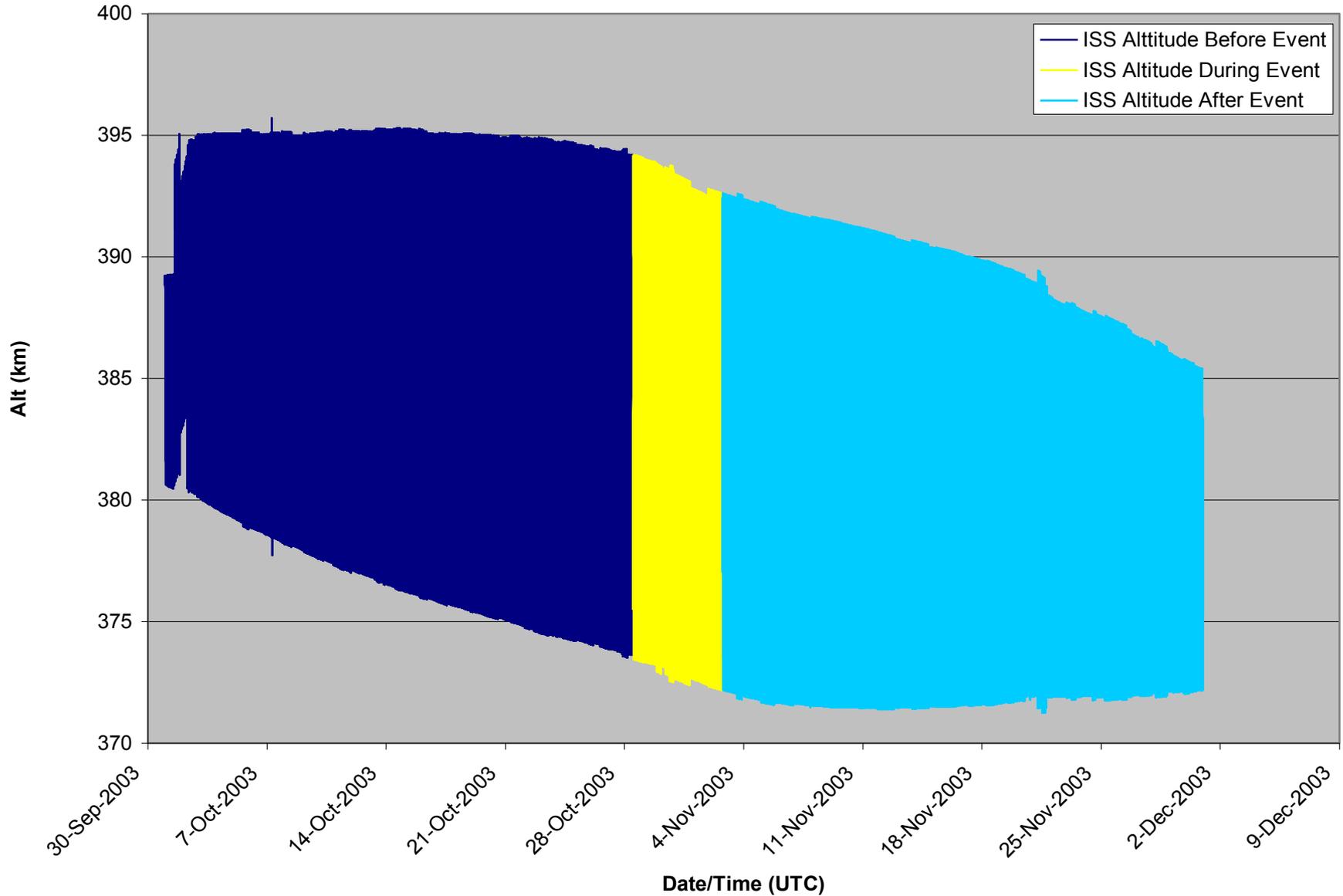


“Toy” Model Satellite Altitude vs Time



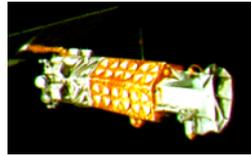
Longer on-orbit lifetime in “cold” thermosphere

ISS Altitude Oct-Nov 2003



Atmospheric Drag = Space Object Positioning Error

**EXPECTED
POSITION**

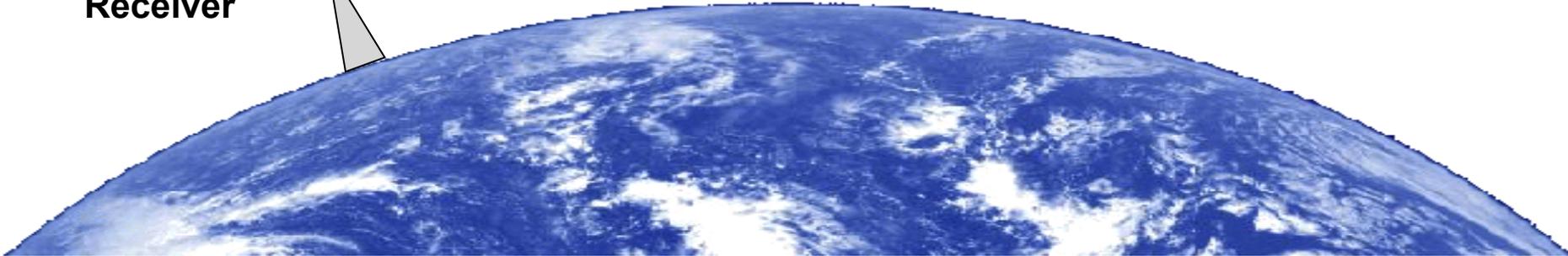
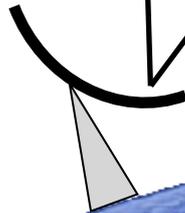


ACTUAL POSITION



Satellite will be some distance below and ahead of its expected position when a ground radar or optical telescope attempts to locate it.

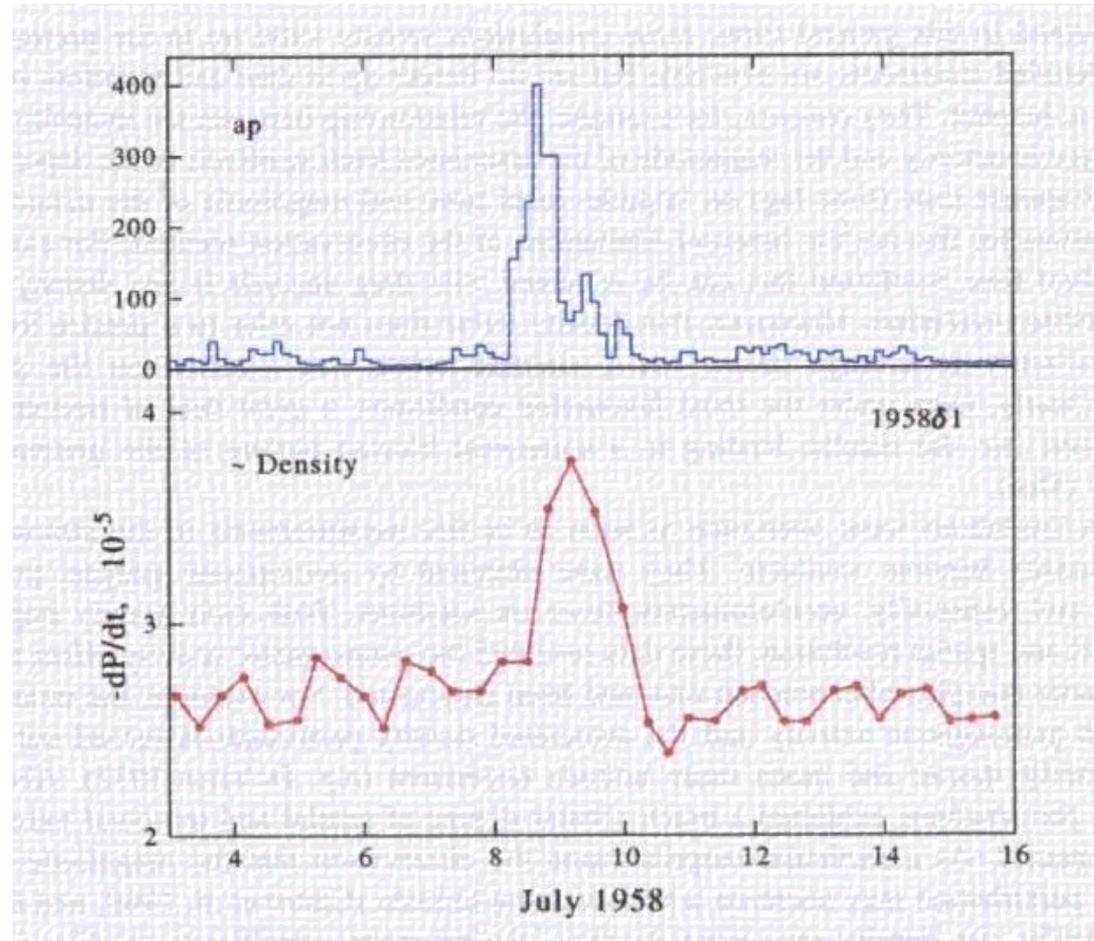
**Radar
Receiver**



First Observation of Satellite Drag Associated with Neutral Density Enhancement

Extreme Geomagnetic Activity Indicated by Auroral Currents

Rate of Orbital Period Change

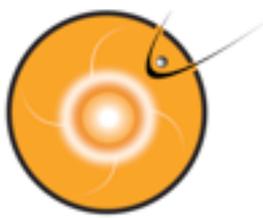


Sputnik Orbital Period Variation vs Day of July 1958

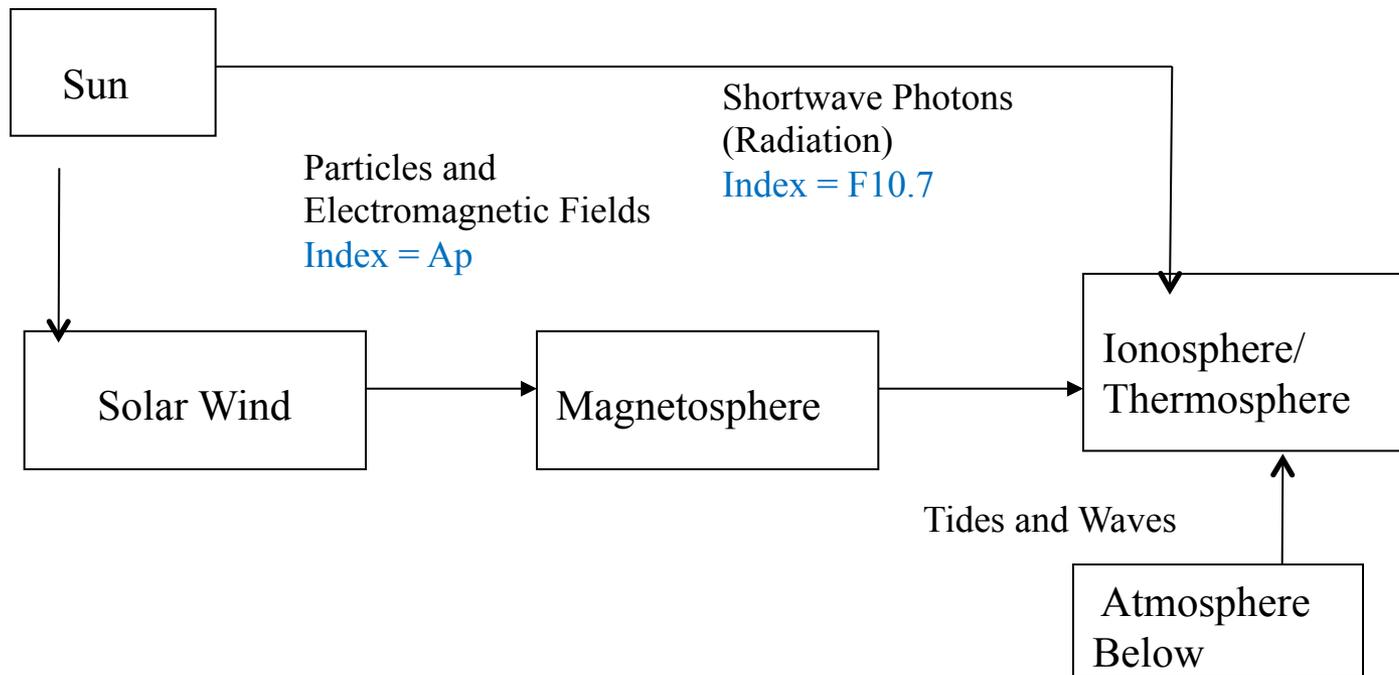
Conjunction Assessment & Collision Avoidance at GSFC

- **Risk Mitigation Maneuvers (RMM) are performed typically ~24 hours prior to the Time of Close Approach (TCA)**
- **Use High Accuracy Satellite Drag Model**
- **Uncertainty due to solar effects still exist**
 - Uncertainties on arrival time and magnitude of Solar Events prior to TCA complicate evaluation in determining if a RMM is warranted or could possibly make matters worse

Courtesy: Bill Guit

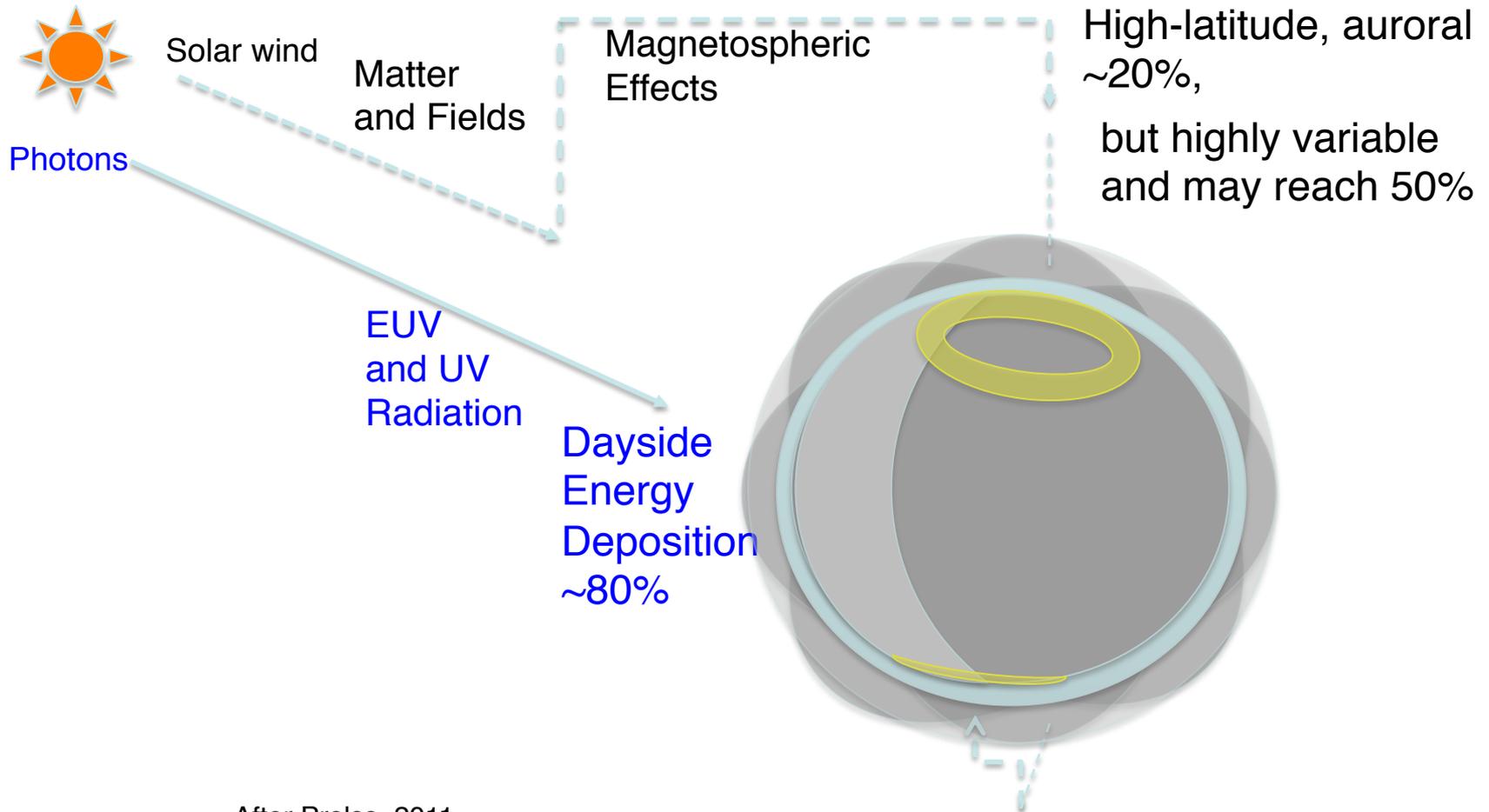


Energy Flow to the Thermosphere

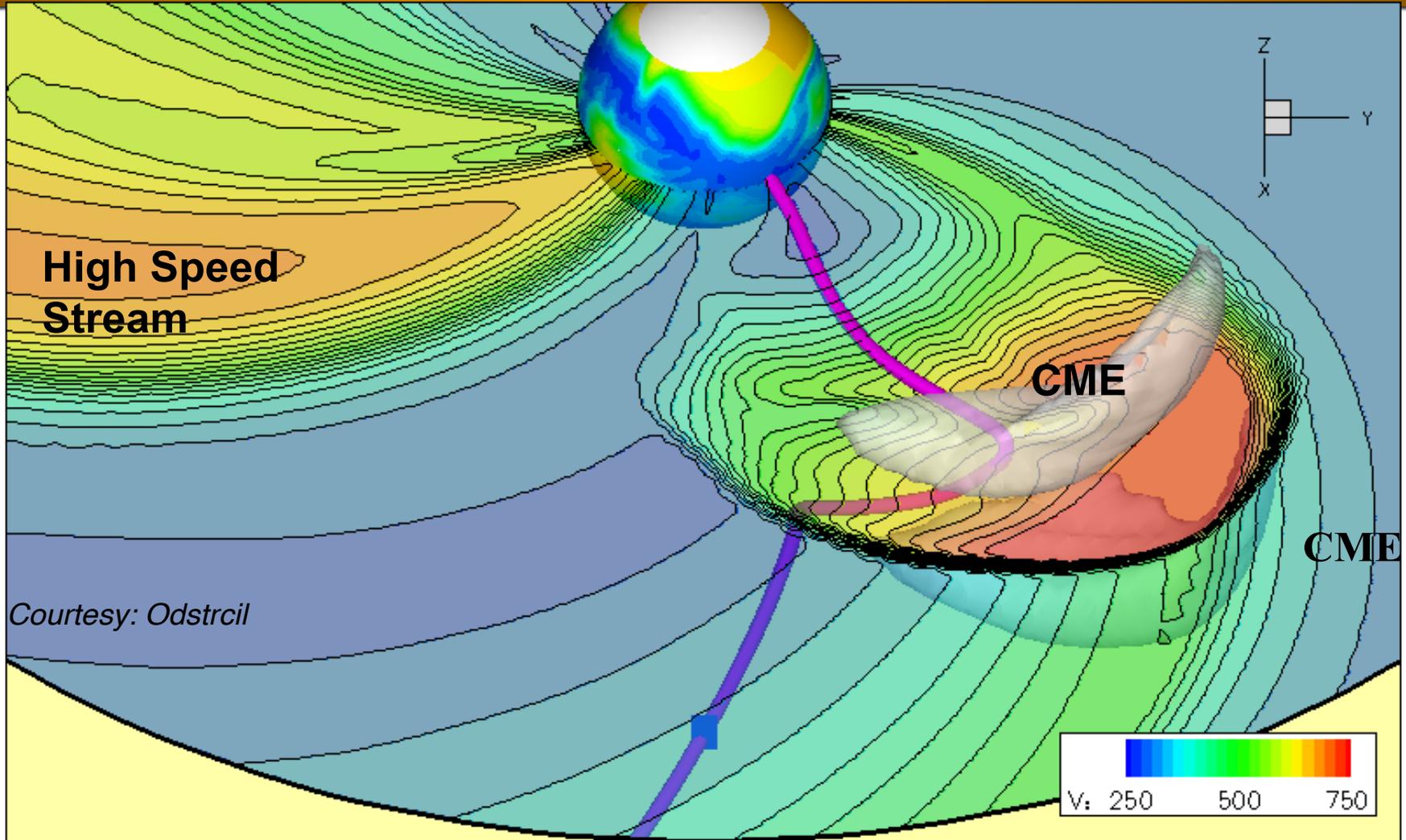


Knipp 2011 Understanding Space Weather and the Physics Behind It

Solar /Solar Wind Energy Deposition

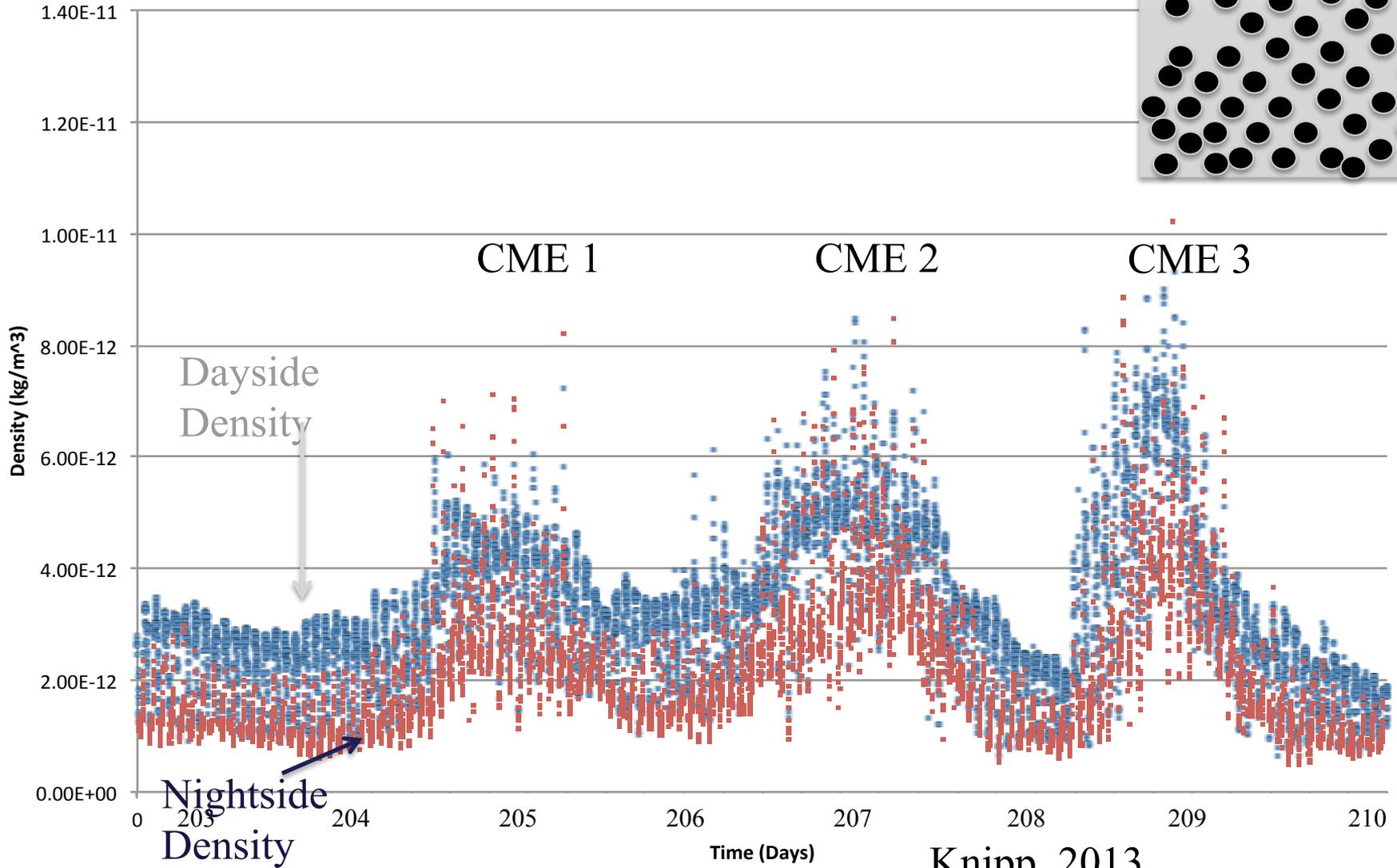
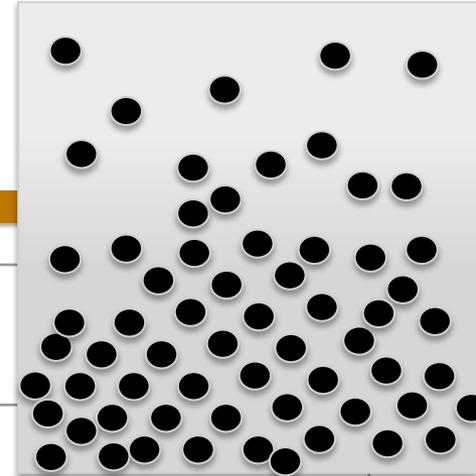


CMEs and HSS Indirectly Heat the Thermosphere



CMEs: extreme short-lived heating, HSSs: Moderate long-lived heating

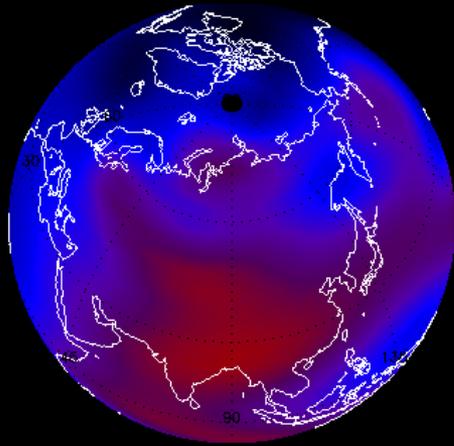
CHAMP Density Extrapolated to 400 km (ng/m^3)



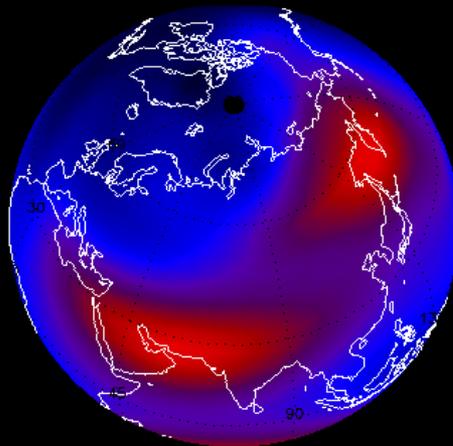
Knipp, 2013

Energy deposition causes atmospheric expansion; Heated molecules and atoms, fighting for more room, migrate upward

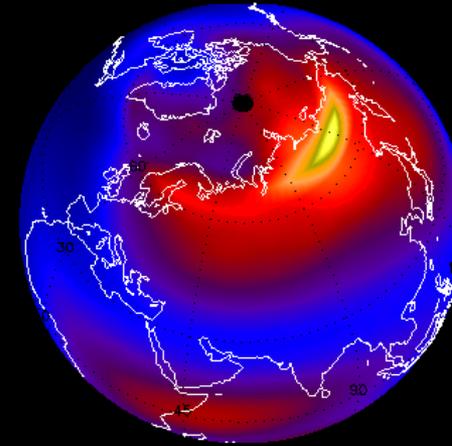
TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 18.00 HEIGHT = 400.00



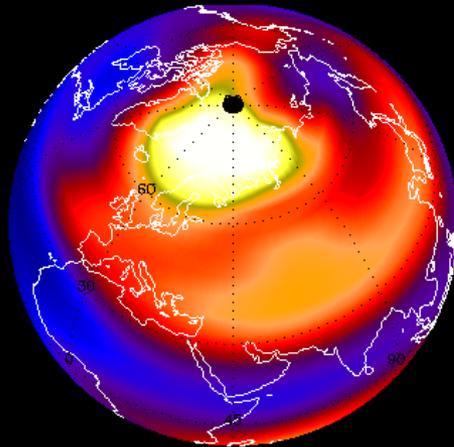
TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 19.00 HEIGHT = 400.00



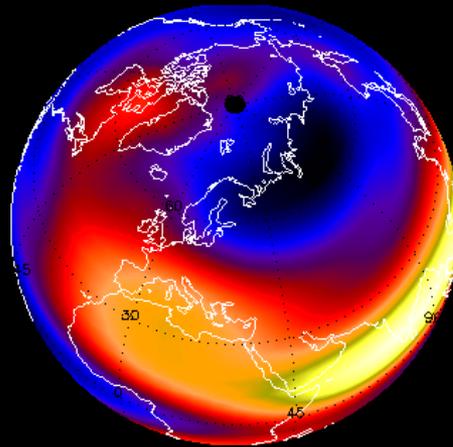
TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 20.00 HEIGHT = 400.00



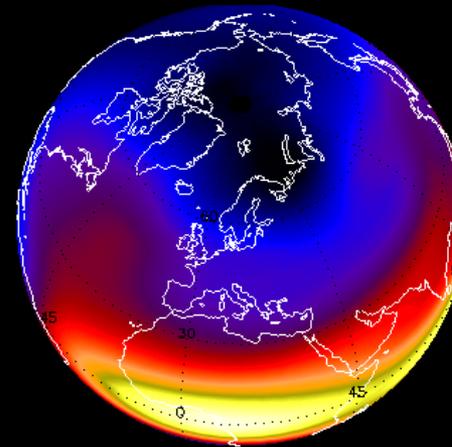
TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 21.00 HEIGHT = 400.00



TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 22.00 HEIGHT = 400.00

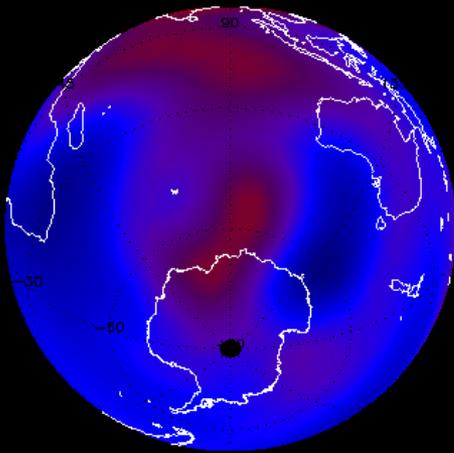


TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 23.00 HEIGHT = 400.00

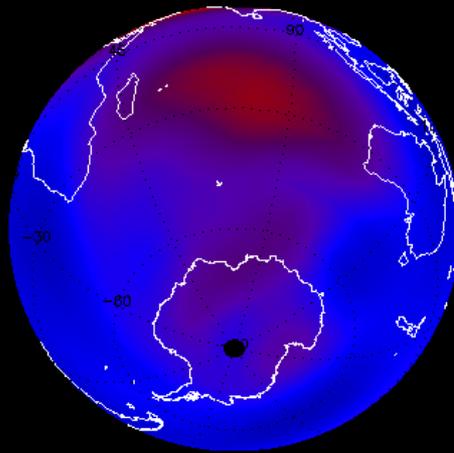


Model output of neutral density change (in %) at 400 km in northern hemisphere during a storm
Atmosphere becomes structured at a fixed altitude. Courtesy of G. Lu, NCAR

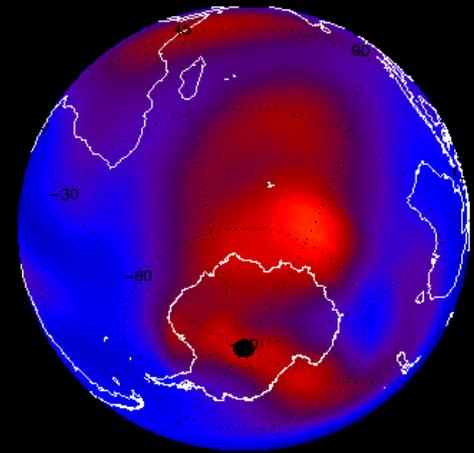
TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 18.00 HEIGHT = 400.00



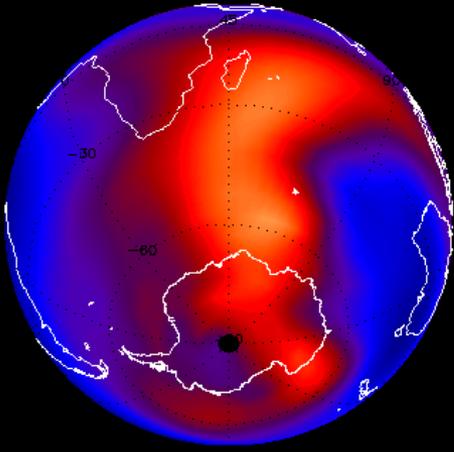
TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 19.00 HEIGHT = 400.00



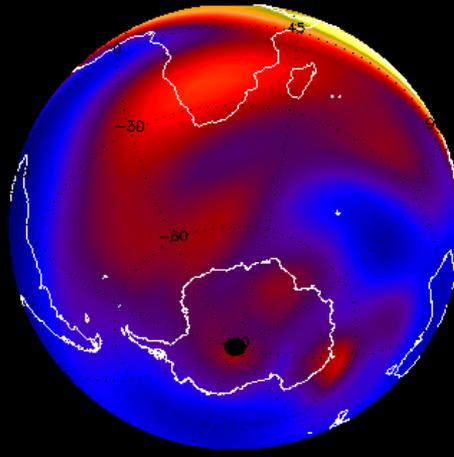
TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
DAY = 314 UT = 20.00 HEIGHT = 400.00



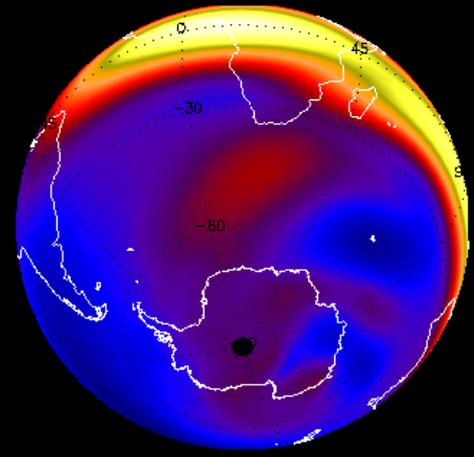
TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
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TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
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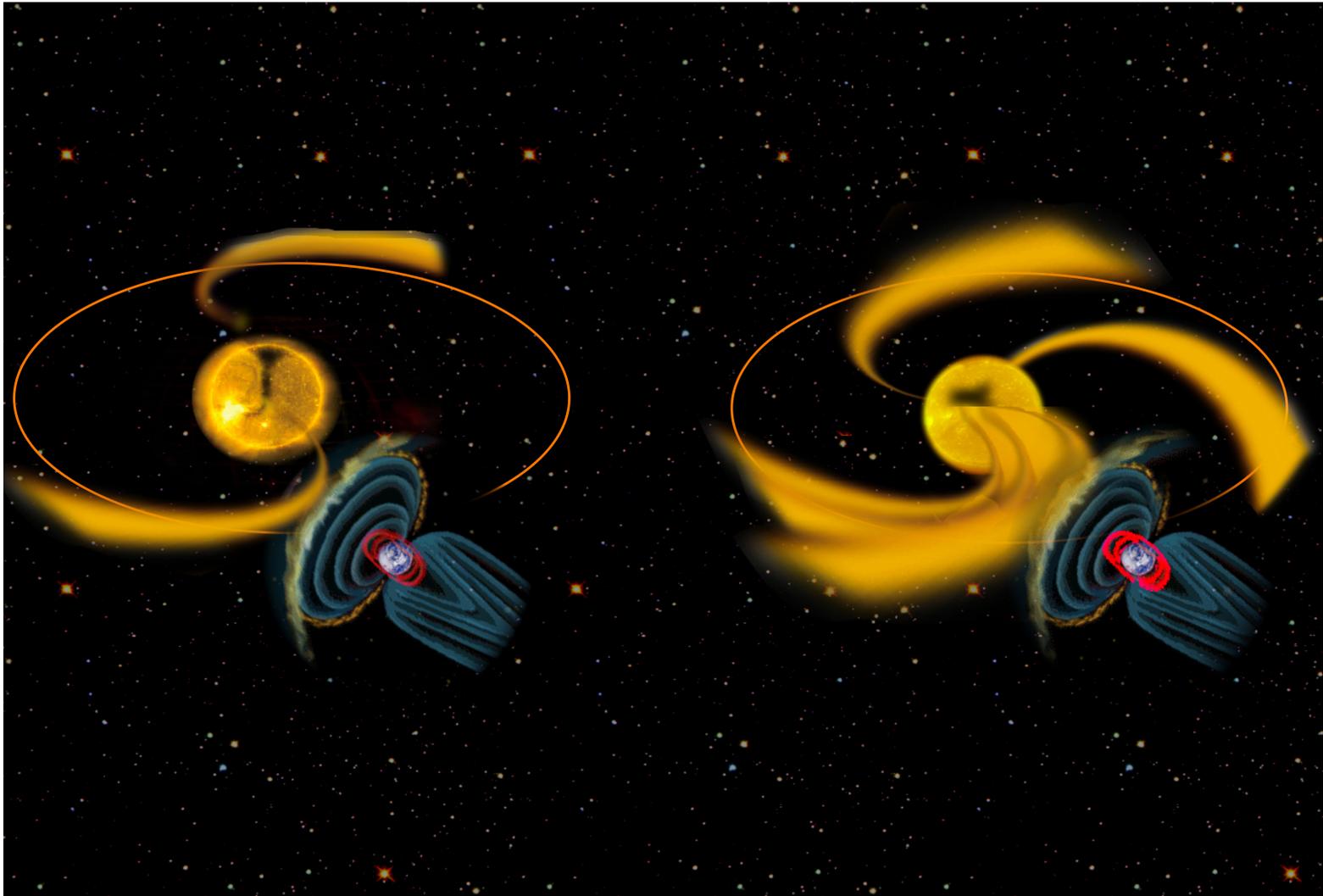


TOTAL DENSITY (O2+O1+N2) (%DIFFS GM/CM3)
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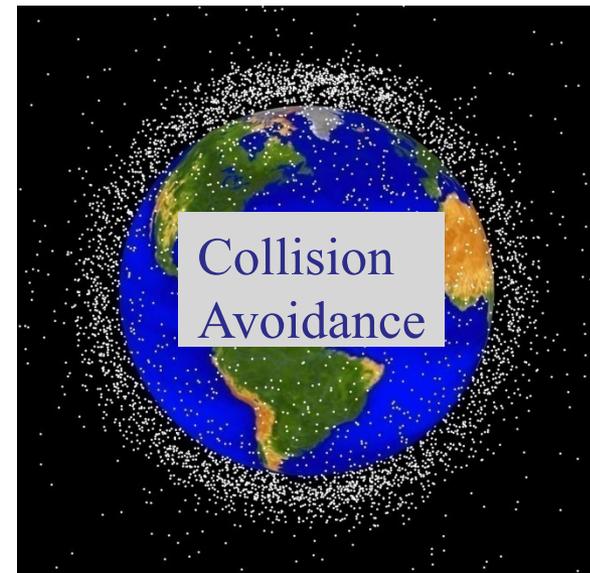
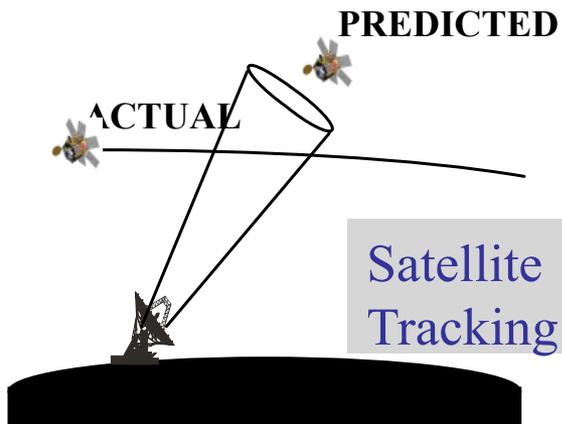
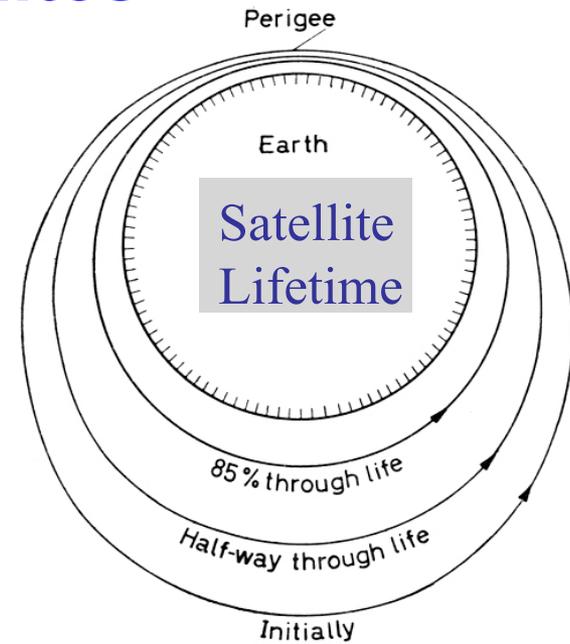
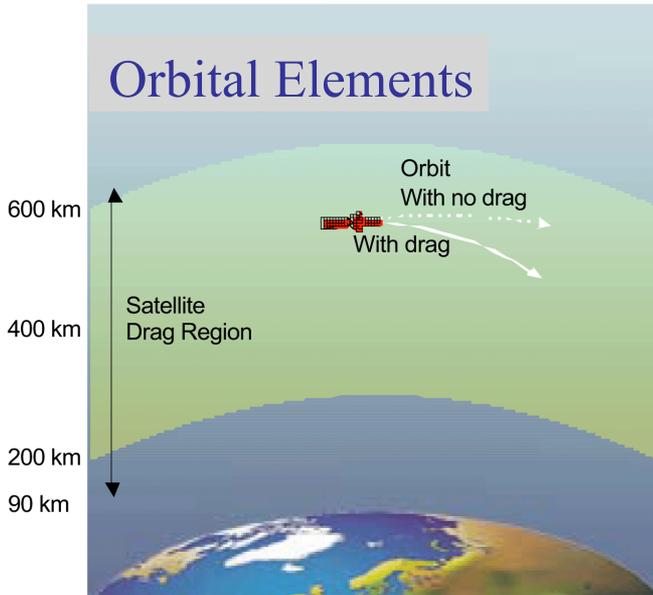
Model output of neutral density change (in %) at 400 km in southern hemisphere during a storm
Atmosphere becomes structured at a fixed altitude. Courtesy of G. Lu, NCAR

High Speed Streams Repetitively Heat the Thermosphere



Gibson et al., 2009

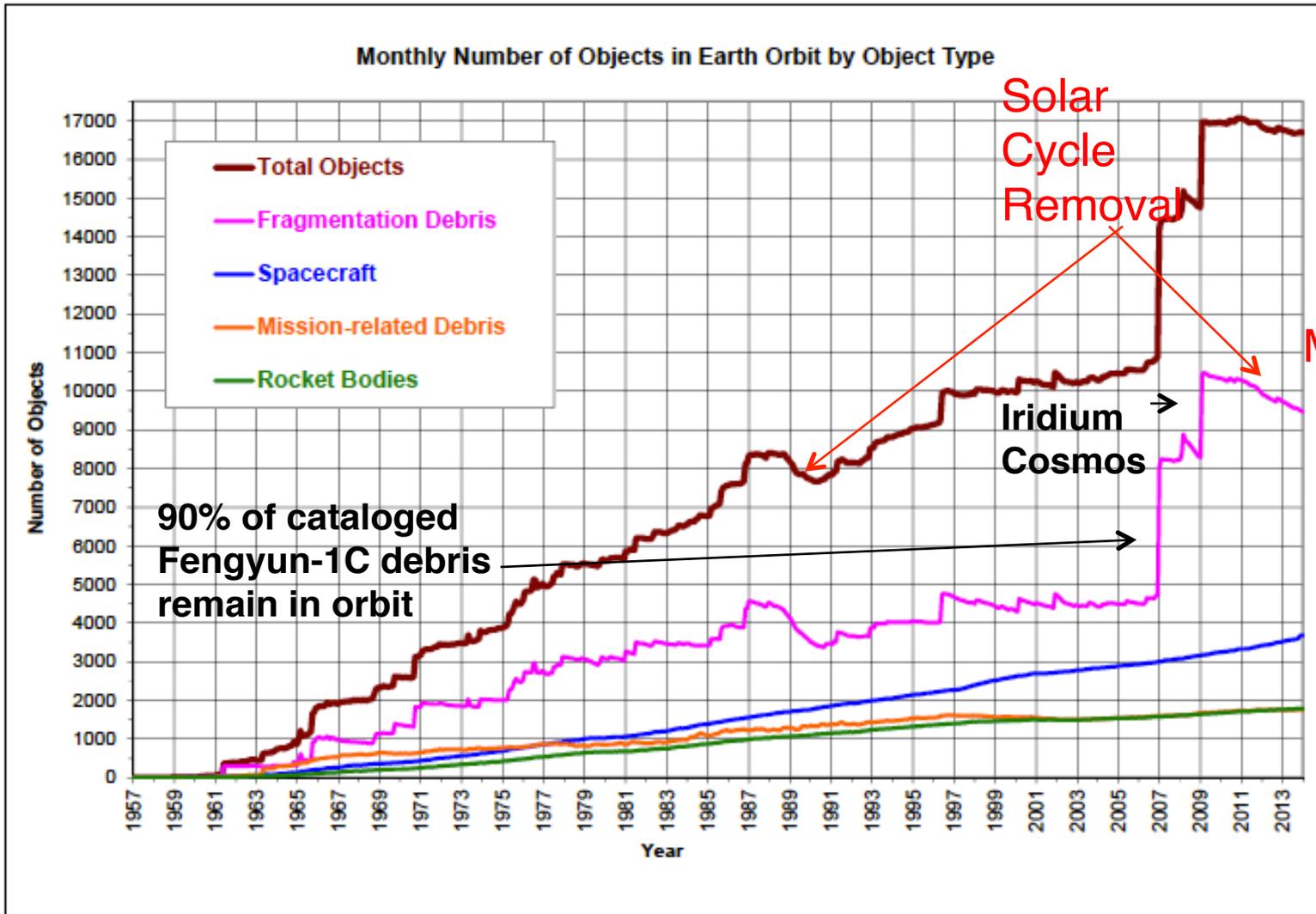
Atmospheric Drag on Satellites



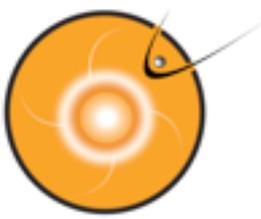
Collision Avoidance

- If a predicted conjunction between orbiting objects and the ISS yields a probability of collision greater than 10^{-4} , official flight rules call for the execution of a collision avoidance maneuver by the ISS. Conjunction volume is: 4km x 50km x 50km box
- During its first 15 years of operations, the ISS successfully conducted 16 collision avoidance maneuvers, and on a separate occasion in 1999 a planned maneuver attempt failed.
- In addition, three incidents arose when insufficient time permitted a collision avoidance maneuver, forcing the crew of the ISS to retreat to the Soyuz return craft where they were prepared to undock from the ISS quickly in the event of a collision.
- In total, the collision avoidance maneuver threshold level has been reached only 20 times for an average of once per year.

Orbital Debris Quarterly News Vol 18, Jan 2014



Monthly Number of Cataloged Objects in Earth Orbit by Object Type: This chart displays a summary of all objects in Earth orbit officially cataloged by the U.S. Space Surveillance Network. "Fragmentation debris" includes satellite breakup debris and anomalous event debris, while "mission-related debris" includes all objects dispensed, separated, or released as part of the planned mission.



Summary



Significant Challenges are posed by satellite drag

Track and identify active payloads and debris

Collision avoidance and re-entry prediction

Attitude Dynamics

Constellation control

“Drag Make-Up” maneuvers to keep satellite in control box

Delayed acquisition of SATCOM links for commanding /data transmission

Mission design and lifetime